BAROCLINIC INSTABILITY

R Grotjahn, University of California, Davis, CA, USA

doi:10.1006/rwas.2002.0076

Grotjahn, R
University of California,
Department of Land, Air, and Water Resources,
Davis, CA 95616-8627, USA

Introduction

0076-P0005 Baroclinic instability refers to a process by which perturbations draw energy from the mean flow potential energy. The conversions of energy are proportional to perturbation heat fluxes in both horizontal and vertical directions. The atmosphere requires heat fluxes to maintain the observed pattern of net radiation (positive in the tropics, negative poleward of 38° N or S on an annual average). A zonal mean meridional circulation, such as a tropical Hadley cell, can generate these heat fluxes. However, in middle latitudes, various factors cause eddies to accomplish the bulk of the heat transport. Baroclinic instability provides a mechanism to explain how these eddies form and evolve while incorporating the necessary heat fluxes. Theoretical models of baroclinic instability can simulate various observed properties of midlatitude eddies, including the dominant length scales, propagation speed, vertical structure, and energetics.

0076-P0010 Baroclinic instability can be viewed as a shear instability. From thermal wind balance, the vertical shear of the zonal wind is proportional to a meridional temperature gradient. The meridional temperature gradient is proportional to the available potential energy (APE) that the baroclinic instability mechanism taps. Another view of baroclinic instability emphasizes interacting potential vorticity (PV) anomalies. Baroclinic instability is usually studied by linearizing the dynamics equations and using eigenvalue or initial value techniques. These alternative views and analysis procedures generally provide complementary means to understand better baroclinic instability.

An Illustrative Model

0076-P0015 An illustrative model provides mathematical relations and archetype solutions for the concepts that follow.

Mathematical Formulation

0076-P0020 The model uses quasi-geostrophic (QG) approximations and nondimensional scaling appropriate for midlatitude frontal cyclones. Potential vorticity has contributions from the interior and from temperature gradients at rigid bottom ($z = 0$) and top ($z = Z_T$) boundaries. In the QG system, PV can be written as:

$$q = \underbrace{\nabla^2\psi}_{\text{RV}} + f_0 + \beta y + \underbrace{\gamma\frac{\partial\psi}{\partial z} + \varepsilon\frac{\partial^2\psi}{\partial z^2}}_{\text{TV}} + \left\{\underbrace{\varepsilon\frac{\partial\psi}{\partial z}\Big]_{z=0} - \varepsilon\frac{\partial\psi}{\partial z}\Big]_{z=Z_T}}_{\text{BPV}}\right\} \quad [1a]$$

where

$$\varepsilon = \frac{f_0^2 L^2}{g\kappa D}, \quad \kappa = D\,\frac{\partial \ln\theta_s}{\partial z}, \quad \text{and} \quad \gamma = \frac{\varepsilon}{\rho}\frac{\partial\rho}{\partial z} + \frac{\partial\varepsilon}{\partial z} \quad [1b]$$

where ψ is the horizontal velocity streamfunction, ρ is density, g is the acceleration due to gravity, and κ is the static stability from the horizontal mean potential temperature. The coordinates are x eastward, y northward, and z upward. Nondimensional length scales are L in the horizontal and D in the vertical. f_0 is the constant part while β is the meridional derivative (approximated as a constant) of the Coriolis parameter.

0076-P0025 An inherent horizontal length scale is the Rossby radius of deformation ($L_R = NHf_0^{-1}$) where N is the Brunt Väisälä frequency ($N^2 = g\kappa D^{-1}$) and $H = RTg^{-1}$ is the scale height (an inherent vertical length scale). Thus, $\varepsilon = (LH)^2(L_R H)^{-2}$ relates the assumed scales L and D to L_R and H.

0076-P0030 Quasi-geostrophic PV includes three distinct parts: absolute vorticity, which includes relative vorticity (RV); "thermal" vorticity (TV); and boundary PV (BPV). Positive PV is associated with an interior trough (in geopotential) and/or a warm surface temperature anomaly.

0076-P0035 When the vorticity and potential temperature conservation equations are combined, one obtains a time-dependent equation for QGPV conservation:

$$\left[\frac{\partial}{\partial t} + U\frac{\partial}{\partial x}\right]\left(\nabla^2\psi + \frac{1}{\rho}\frac{\partial}{\partial z}\left(\rho\varepsilon\frac{\partial\psi}{\partial z}\right)\right) + \frac{\partial Q}{\partial y}\frac{\partial\psi}{\partial x} = 0 \quad [2a]$$

with boundary conditions at the bottom and top

$$\left[\frac{\partial}{\partial t} + U\frac{\partial}{\partial x}\right]\left(\frac{\partial\psi}{\partial z}\right) - \frac{\partial U}{\partial z}\frac{\partial\psi}{\partial x} = 0 \quad \text{at} \quad z = 0, Z_T \quad [2b]$$

"Basic state" variables are specified: U (independent of x) is zonal wind, and Q is the interior part of the QGPV; meridional and vertical velocities are zero. One can solve eqn [2] as an initial value problem by specifying an initial streamfunction or potential vorticity.

0076-P0040 An eigenvalue problem can also be formulated from eqn [2]. A common approach assumes time and space dependence in the form:

$$\psi(x, y, z, t) = \mathrm{Re}\{\phi(y, z)\exp[\mathrm{i}k(x - ct)]\} \quad [3]$$

for the "perturbation" streamfunction being sought. This solution has zonal wavenumber k and complex phase speed c. The growth rate is given by $k\mathrm{Im}\{c\}$. If U has no meridional variation, then one can assume a wavelike y dependence too: $\exp(\mathrm{i}ly)$. When wavenumber $l = k$, the solution is a "square wave." Perturbation velocities are defined as $u = -\partial\psi/\partial y$ and $v = -\partial\psi/\partial x$.

0076-P0045 Additional simplifying approximations are often made. A particularly simple form, commonly labelled the "Eady model," was described by E. T. Eady in 1949. The Eady model assumes wavelike meridional structure, $\partial Q/\partial y = 0$, $U = z$, incompressibility (ρ = constant), and $\varepsilon = 1$. Then eqn [2a] is reduced simply to solving $q' = 0$ in the interior where the prime denotes the "perturbation" sought. The Eady eigenvalue problem can be solved analytically to yield a pair of normal modes, one growing, and one decaying, for scaled wavenumber $\alpha < \sim 2.4$. The scaled wavenumber:

$$\alpha = \left\{(k^2 + l^2)\varepsilon^{-1}\right\}^{1/2} \quad [4]$$

is proportional to absolute wavenumber and static stability.

0076-P0050 Equations for perturbation kinetic energy, K_e and available potential energy, A_e are:

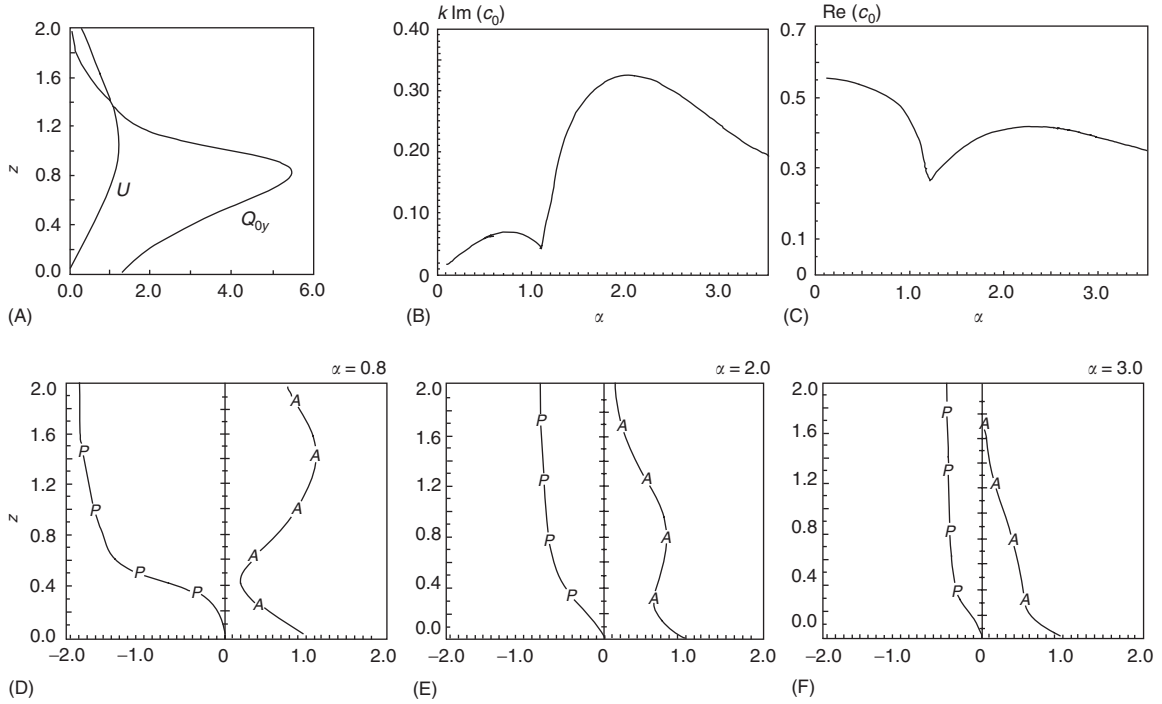
$$\frac{\partial A_e}{\partial t} \equiv \frac{\partial}{\partial t}\int\int\int \frac{\rho_S}{2}\left\{\varepsilon\left(\frac{\partial\psi}{\partial z}\right)^2\right\}\mathrm{d}x\,\mathrm{d}y\,\mathrm{d}z = \underbrace{\int\int\int \varepsilon\rho_S\frac{\partial U}{\partial z}\frac{\partial\psi}{\partial x}\frac{\partial\psi}{\partial z}\,\mathrm{d}x\,\mathrm{d}y\,\mathrm{d}z}_{(A_z\to A_e)} - \underbrace{\int\int\int \rho_S w\frac{\partial\psi}{\partial z}\,\mathrm{d}x\,\mathrm{d}y\,\mathrm{d}z}_{(K_e\to A_e)} \quad [5a]$$

$$\frac{\partial K_e}{\partial t} \equiv \frac{\partial}{\partial t}\int\int\int \frac{\rho_S}{2}\left\{\left(\frac{\partial\psi}{\partial y}\right)^2 + \left(\frac{\partial\psi}{\partial x}\right)^2\right\}\mathrm{d}x\,\mathrm{d}y\,\mathrm{d}z = \underbrace{-\int\int\int \rho_S U\left\{\frac{\partial\psi}{\partial x}\frac{\partial\psi}{\partial y}\right\}_y\mathrm{d}x\,\mathrm{d}y\,\mathrm{d}z}_{(K_z\to K_e)} + \underbrace{\int\int\int \rho_S w\frac{\partial\psi}{\partial z}\,\mathrm{d}x\,\mathrm{d}y\,\mathrm{d}z}_{(A_e\to K_e)} \quad [5b]$$

The volume integrals are over a closed domain. In the QG system, $\partial\psi/\partial z$ is proportional to potential temperature θ, making the first term on the right-hand side of eqn [5a] proportional to a meridional heat flux, while the second term is proportional to a vertical heat flux. The specified vertical shear, $\partial U/\partial z$, is proportional to the available potential energy, A_z of the basic state and is the energy upon which the baroclinic instability mechanism feeds. The first term on the right-hand side of eqn [5b] is a barotropic energy conversion. The barotropic conversion is proportional to the divergence of eddy momentum flux and also draws energy from the mean flow. The second term on the right-hand side of eqns [5a] and [5b] is the same, but with opposite sign indicating a conversion between A_e and K_e.

Example Solutions

0076-P0055 This QG eigenmodel of baroclinic instability is applicable to the midlatitudes. In these regions zonal flow increases with height reaching a maximum near the tropopause. **Figure 1A** is a representative nondimensional profile of U where the tropopause is at nondimensional $z = 1.0$. The growth rate and phase speed spectra, along with the (growing normal mode) eigenfunction structures for different k are also shown in **Figure 1**. The growth rate has maximum value at a specific value of α. The vertical structure tends to have relative maxima at the surface and near the tropopause, but it becomes progressively more bottom-



0076-F0001 **Figure 1** Quasigeostrophic eigenanalysis. (A) Specified zonal wind U, and meridional gradient of interior potential vorticity Q_{0y} versus scaled height. $z = 1$ is 10 km. (B) Growth rate and (C) phase speed versus absolute wavenumber α. (D)–(F) Amplitude A, and phase P, for the growing normal mode for $\alpha = 0.8$, $\alpha = 2.0$, and $\alpha = 3.0$, respectively. All three modes tilt westward (upstream) with increasing height. Dimensional wavelengths depend upon scaling assumptions, but reasonable choices imply that $\alpha = 0.8$, $\alpha = 2.0$, and $\alpha = 3.0$ correspond to 11.0, 4.4, and $\sim 2.9 \times 10^3$ km wavelengths, respectively. (Zonal and meridional scales are set equal.) The same scaling implies phase speed of 9 m s^{-1} and doubling time of $\sim$1.2 days for $\alpha = 2.0$. (Adapted with permission from Grotjahn R (1980) *Journal of Atmospheric Sciences* 37: 2396–2406 by permission.)

trapped for shorter waves. The phase varies such that unstable modes tilt upstream with height, i.e., against the mean flow shear. Other solutions to eqn [2], labeled continuum modes, are relevant to nonmodal growth.

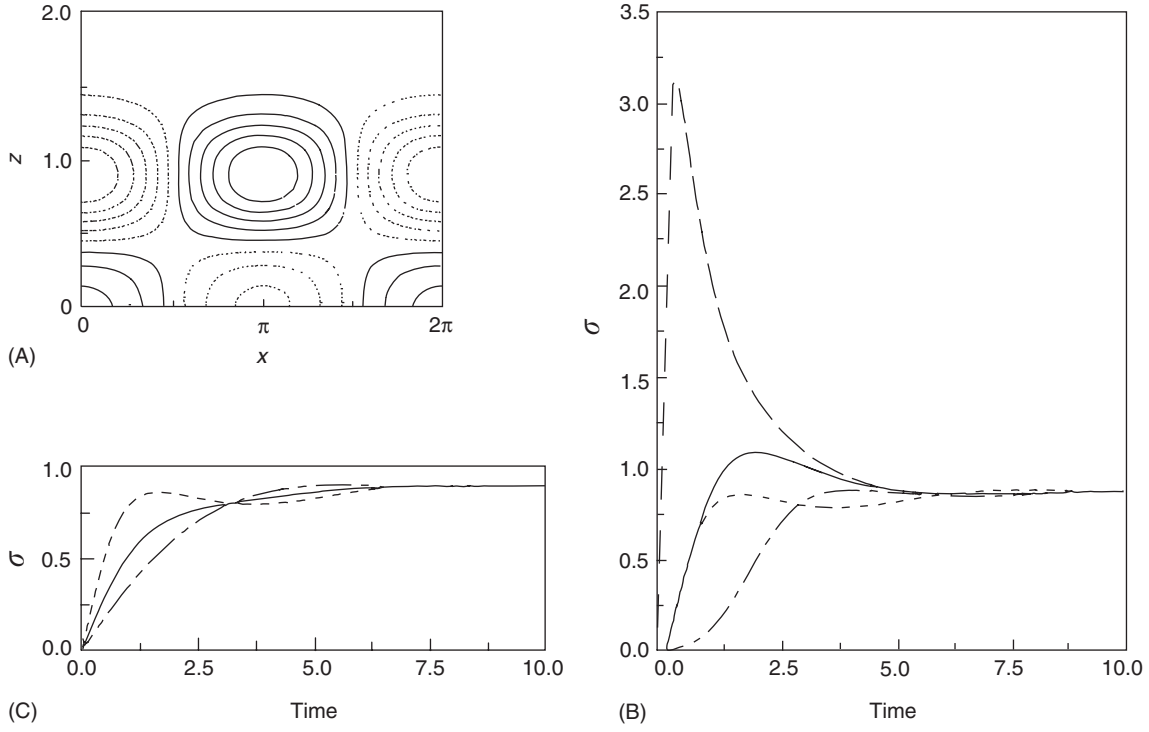
0076-P0060 For shorter waves the lower maximum tends to dominate (when compressibility is included) and the solution decays rapidly away from the boundary. For longer waves the tropopause level maximum tends to dominate (**Figure 1C**). Eady model normal modes have interior $q' = 0$; from eqn [1]: the Laplacian increases as k^2, requiring a rapid change with height for short waves to make the thermal term comparable (this leads to boundary trapping of the solutions). For longer waves, the Laplacian becomes small and the vertical structure is more evenly spread in the vertical, hence these modes are "deeper."

0076-P0065 Typical geopotential patterns observed prior to frontal cyclone development have separate surface and upper troposphere troughs, each equivalent barotropic (vertical trough axis), with the upper-level trough more prominent. A crude simulation of that initial state is used to generate solutions shown in **Figure 2**. Time-series of the growth rates of several quantities are tracked over several days. The time-series include potential enstrophy $(H(q')^2)$ and total energy $(TE = A_e + K_e)$ integrated over the whole domain. The solutions approach asymptotically to the most unstable normal-mode growth rate as that eigenmode emerges to dominate the solution. The growth rate has transient peak values that can exceed the asymptotic (normal-mode) value.

Classical View

Baroclinic instability draws upon the APE of the 0076-P0070
environment in which an eddy sits. Since APE is related to a horizontal temperature gradient, which is in turn related to the vertical shear, it can be viewed as a type of shear instability. One advantage of doing so is to make comparisons with barotropic instability, which draws energy from the horizontal shear. This view provides a link to the eddy fluxes that are observed and necessary for each conversion.

As demonstrated in eqn [5], heat fluxes are neces- 0076-P0075
sary to have a baroclinic energy conversion. Horizontal heat fluxes imply that the temperature and mass (here ψ) fields are offset. The offset implies that the trough and ridge axes tilt upstream with elevation.



0076-F0002 **Figure 2** Initial value calculation. (A) Zonal cross-section of initial streamfunction, dashed contours used for negative values. (B) Time-series of growth rates for domain average potential enstrophy (solid line) and its components: RV^2 (short dashed line), TV^2 (dot-dashed line), and BPV^2 (long dashed line). Growth rates approach asymptotically the most unstable normal mode rate for this wavenumber $\alpha = 2.0$. (C) Similar to (B) except for total energy (solid line), kinetic energy (short dashed line), and available potential energy (dot-dashed line).

0076-P0080 The QG formulation above is adiabatic, so individual parcels conserve their potential temperature (θ) over time. For unstable modes, the horizontal and vertical eddy heat fluxes must distort the θ field over time, as suggested schematically in **Figure 3**. An isentropic (θ) surface is drawn in three-dimensional perspective; it curves up and over colder areas and dips down over warmer areas. Prior to eddy development, the isentropic surface did not vary in the x direction and had a shape like its intersection with the wall at $x = 0$. The isentropic surface is distorted by flow around the high- and low-pressure centers and representative cold (C) and warm (W) trajectories are also drawn. When these trajectories are projected onto the $x = 0$ wall, they appear to cross the initial zonal mean isentrope and have a slope that is typically half the slope of the mean isentrope. In fact, they are changing the zonal mean of the isentrope to become more horizontal, thereby reducing the horizontal temperature gradient and thus reducing A_z. In this classical view, A_z is reduced while A_e is increased by increasing the zonal undulations of the isentropic surface. Another aspect is that colder air is sinking while warmer air is rising, a process that lowers the center of mass and thus converts A_e into K_e. To lower the center

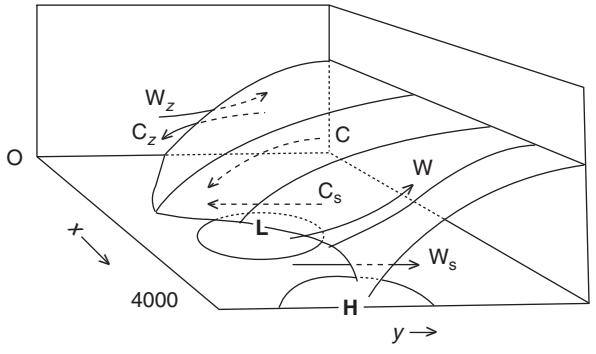


Figure 3 Schematic diagram showing distortion of an isentropic surface by a baroclinically amplifying frontal cyclone. Dotted lines are used for objects underneath the three-dimensional isentropic surface. Surface high, H, and low, L, are marked together with two representative contours of surface pressure. Trajectories of representative parcels are shown in warm air, W, and cold air, C. Subscript s denotes projection onto the bottom surface, while z denotes projection onto the meridional plane (where $x = 0$). The trajectories do not cross the isentropic surface but distort it. Initially the isentropic surface had negligible variation with x and looked like the current pattern at $x = 0$. The projections W_z and C_z seem to cross the initial isentropic surface but in fact are flattening it (which reduces A_z). Rising air is warm while sinking air is cold, which lowers the center of mass, converting A_e into K_e. 0076-F0003

of mass, the parcel paths must have the vertical component indicated, but they must also be less than the slope of the mean isentropes for instability to occur.

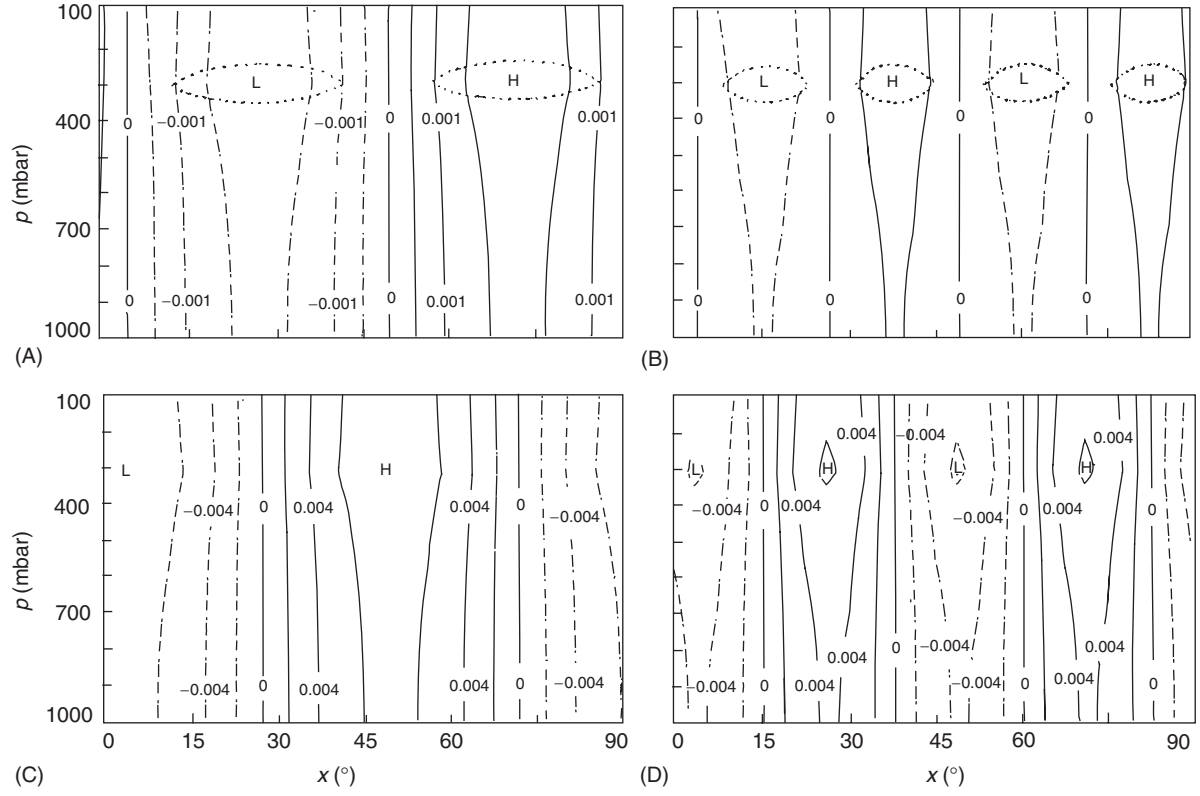
0076-P0085 The classical view can incorporate latent heat release as follows. The bulk of the precipitation in a developing cyclone forms in the warm-air sector of the storm. The release of latent heat further depresses the isentropic surfaces where there is poleward motion, implying additional conversion of A_z into A_e and K_e.

Potential Vorticity View

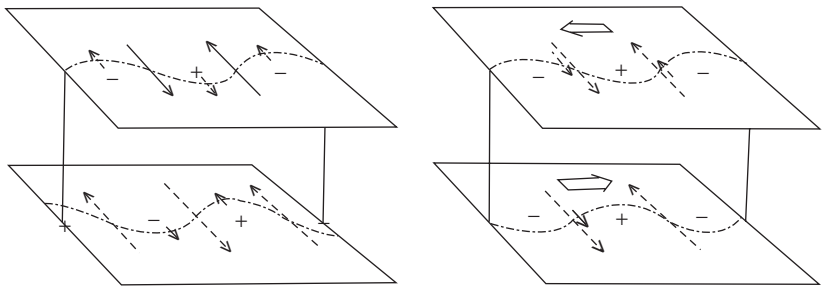
0076-P0090 The potential vorticity view of instability tracks how two or more PV anomalies interact in a way that causes growth of the PV anomalies. Potential vorticity is a fundamental conserved quantity for adiabatic motions. The illustrative model is designed around QGPV conservation.

0076-P0095 A PV pattern has an associated streamfunction and horizontal wind field. In general, eqn [1] implies that PV emphasizes smaller-scale variations than the streamfunction field. Inverting eqn [1] obtains broad patterns of ψ associated with isolated packets of q. An illustrative example is shown in **Figure 4** for PV anomalies in the upper troposphere. The associated winds are displaced from the PV anomaly center by $\frac{1}{4}$ wavelength ($\sim$1000 km). A similar diagram can be constructed for a PV anomaly associated with a surface temperature gradient.

Potential vorticity anomalies are created by flow 0076-P0100 across PV contours. **Figure 5** illustrates how two sinusoidal PV anomalies can amplify each other. The PV gradient is reversed between the two levels, increasing with y at upper levels and decreasing with y at the surface. This pattern is consistent with upper tropospheric PV and the surface temperature gradient, respectively. (Recall that q is positive for lower geopotential height or warmer surface temperature.) The associated winds cause propagation by distorting the PV pattern. However, the meridional wind associated with a PV anomaly is in quadrature with that anomaly so the PV cannot amplify itself. Growth is described simply as advection at the PV extrema that further amplifies the PV pattern. Since the associated winds extend beyond the elevation of the PV anomaly,



0076-F0004 **Figure 4** Quasi-geostrophic potential vorticity inversion on a sphere. Zonal cross-sections are shown over 90° longitude. Potential vorticity anomalies are specified and centered at 45° N; (A) and (B) use dotted lines to indicate PV contours 0.5 and -0.5 nondimensional units for wavenumbers 4 and 8, respectively. Contours are streamfunction, ψ, derived from inverting the PV relation of eqn [1]. Lows ($\psi < 0$) correspond with PV > 0 maxima, and vice versa. (C) and (D) show the meridional winds associated with ψ in (A) and (B), respectively. Dashed contours denote negative values. The contour interval in (A) and (B) is 0.005; in (C) and (D) it is 0.002. Longer PV anomalies have deeper and stronger streamfunctions.



0076-F0005 **Figure 5** Baroclinic instability from interacting PV anomalies at two levels. A representative PV contour (dot-dashed line) is drawn at each level. The offset is (A) $\frac{1}{4}$ wavelength and (B) $\frac{1}{2}$ wavelength. A typical wavelength might be 4×10^3 km. Each anomaly has an associated wind component parallel to the PV gradient; dashed arrows are winds from the lower PV anomaly, while solid arrows are from the upper anomaly. The winds from each anomaly advect the associated anomaly. In (A) each PV anomaly has a wind component that amplifies the undulation in the other anomaly, thereby causing growth. In (B) each PV anomaly has a wind component that augments the propagation in the manner indicated by the broad arrows; this causes the anomalies to migrate to a phaselike diagram (A).

there can be interaction with a second PV anomaly at another level. When the second PV anomaly is offset from the first, as in **Figure 5A**, the associated winds amplify the first anomaly.

0076-P0105 This mechanism also explains how developing cyclones maintain a preferred tilt (i.e., become "phase locked"). The lower anomaly is shifted horizontally to the right in **Figure 5B** so that upper and lower anomalies are 180 degrees out of phase. The two PV anomalies no longer amplify each other's PV anomalies (shutting off the instability mechanism). Furthermore, the two anomalies reinforce the velocities midway between their positive and negative extremes, thereby enhancing the propagation at each level; but the propagation is in opposite directions at each level, thereby reducing the phase shift to reestablish the pattern in **Figure 5A**. As with the classical view, normal modes are a special case where this phase locking is optimized.

0076-P0110 The PV view provides theoretical weight to a classic description of how cyclones develop: an upper-level trough (PV anomaly) approaches a low-level baroclinic zone (another PV anomaly), then growth commences. This paradigm is commonly labeled "type B" cyclogenesis.

0076-P0115 Observations show independent troughs at the upper and lower troposphere prior to development with the upper approaching the lower. Neither trough has upstream tilt initially; such tilt emerges only after the two become favorably aligned and growth has commenced.

0076-P0120 A necessary condition for instability is that the across-flow mean gradient of PV changes sign within the domain. In the illustrative model above, $\beta > 0$ meaning $Q_y > 0$ in the interior, and the surface temperature gradient $(\mathrm{d}\theta/\mathrm{d}y < 0)$ implies $Q_y < 0$ at the ground. In the Eady model $Q_y = 0$ everywhere in the interior, so the normal-mode instability comes from BPV having opposite sign at top and bottom boundaries.

A necessary condition for instability is that a 0076-P0125
steering level, where $U = \mathrm{Re}\{c\}$, lies within the domain. A supportive kinematic argument is that air parcels remain with the system (rather than blow through it or be left behind) and are more easily mixed laterally. For really long waves, strong retrogressive motion caused by the β term leads to a different class of unstable eigenmodes for $\alpha < \sim 1.1$ (note cusp in **Figure 1B**) than for larger α.

Normal Modes

Normal modes are physically meaningful eigenfunc- 0076-P0130
tions. As in the illustrative model, the equations are linearized about a specified basic state and perturbation solutions are sought. Most commonly, the time and one or more space dependencies are assumed. By assuming a form like that in eqn [3], unstable solutions grow exponentially. Models that are simple enough may be solved analytically. More commonly, the eigenvalue problem is solved numerically.

Normal modes are consistent with many observed 0076-P0135
features:

1. Unstable modes tend to be lined up along the jet axis (if present) in the mean flow.
2. The most unstable wavelength is similar to the observed median size. The normal mode scale can be manipulated by varying the choices made for nondimensional parameters, but is on the order of 4500 km.
3. Solutions tend to develop similar zonal and meridional lengths, the latter responding to the width of the jet that provides one natural scale in the model.

Other properties (such as static stability) also influence the length scales.

4. The vertical structure of the most unstable modes tends to have relative maxima at the surface and upper troposphere.
5. In growing normal modes the temperature lags the mass field (typically by 20–50 degrees of phase for the most unstable mode). Two consequences are:
   - First, the mass field must tilt upstream with height. There is typically $\frac{1}{4}$ to $\frac{1}{2}$ wavelength (1–2×10^3 km) between the trough location at the surface and at tropopause level.
   - Second, the lag allows across-flow heat fluxes down the temperature gradient, as expected from eqn [5a]. In the Eady model the heat flux is uniform with height. Model improvements, most notably compressiblity, can emphasize the eddy heat flux in the lower troposphere (where observations find it most prominent).
6. The rate of propagation is $\sim$10–20 m s^{-1}: slower than jet stream level winds, but faster than (zonal average) surface winds. The steering level is defined as the level where the propagation speed of the storm is equal to the wind component along the storm's track. The steering level for the most unstable normal modes is typically between 700 and 500 mbar, depending on the assumptions made. For shorter waves, the steering level is closer to the surface, and these modes move more slowly. Longer waves respond to competing effects: they have greater upper-level amplitude (where U is faster) but greater sensitivity to β (which enhances retrograde motion).
7. The rate of growth is similar to but slower than that of observed cyclones. Observed doubling times are typically 1–2 days at upper levels.
8. Instability is inversely proportional to static stability. For example, the peak growth rate depends on α ($= 2.0$ in **Figure 1**). From eqns [1b]and [4], α is proportional to static stability κ. Hence, smaller κ places the most unstable peak at larger k, making the growth rate (kIm{c}) larger. Kinematically, vertical motion needed in eqn [5] becomes easier for smaller κ.

0076-P0140 The fact that normal modes have fixed tilt is not necessarily unrealistic. Observations of the vorticity equation terms support an approximately fixed structure for a developing low because the divergence term opposes the horizontal advection at upper levels but reinforces the horizontal advection at low levels. The normal modes (**Figure 1**) are special structures where the net advection is exactly uniform throughout the depth of the fluid. Tracking observed troughs over time shows some evidence for a period of fixed tilt during growth, though the 12 h interval between upper air observations makes the measurement difficult. The vorticity equation also illustrates instability by which the divergence term has positive vorticity tendency at a trough where vorticity is a maximum, thus amplifying the peak vorticity (and vice versa for ridges).

0076-P0145 In addition to the normal modes, the eigenfunctions include a class of solutions called "continuum" modes. For an adiabatic model continuum modes have equivalent barotropic structure (no tilt), making them neutral. In the Eady model, continuum modes have zero PV at all levels except at the critical level, where their amplitude has a "kink." Continuum modes play a role in nonmodal growth.

Nonmodal Growth

0076-P0150 Nonmodal growth is seen when solving initial value problems such as eqn [2]. The formulation can be linear, as in eqn [2], or nonlinear. This approach is more general than eigenanalysis, since the time dependence is not assumed.

0076-P0155 The solution at any time can be decomposed into a combination of eigenfunctions. For an arbitrary initial state, continuum and normal modes are present. These modes move at differing speeds. In a linear formulation the modes operate independently; as modes disperse, positive and negative reinforcement varies. The interference between modes decays algebraically asymptotically. However, for some initial conditions it is possible to have sizable growth over a limited time period.

0076-P0160 For the Eady model, analytic solutions can be found that illustrate the process. Using an initial condition with upstream tilt ($\phi \sim \exp(imz)$ in eqn [3], where $m > 0$) yields solutions with normal mode and algebraic parts. The algebraic part has time dependence proportional to $\{(m - kt)^2 + \alpha^2\}^{-1}$ and $\exp\{i(m - kt)z\}$. The amplitude increases as the tilt becomes more vertical until $t = mk^{-1}$. After that, the wave tilts downstream and decays.

0076-P0165 Initial upstream tilt becoming more vertical with time has led to an expectation that RV increases at the expense of TV while interior PV remains conserved. However, exceptions can be found where large nonmodal growth occurs (in H) as upstream tilt "develops" from an initial state with no tilt. The explanation lies in a rough cancellation between RV and TV, leaving the BPV evolution to dictate rapid growth in H.

0076-P0170 A robust interpretation of nonmodal growth is progressively more favorable superposition of constituent modes. Continuum modes, having mainly upper-

level amplitude, tend to move quickly, while modes with mainly lower-level amplitude move slowly. Decomposition into eigenmodes of an initial state with upstream tilt finds faster continuum modes located upstream of slower continuum modes. Over time, the modes become more favorably lined up; the tilt becomes more vertical and the total amplitude increases. **Figure 6** illustrates the process.

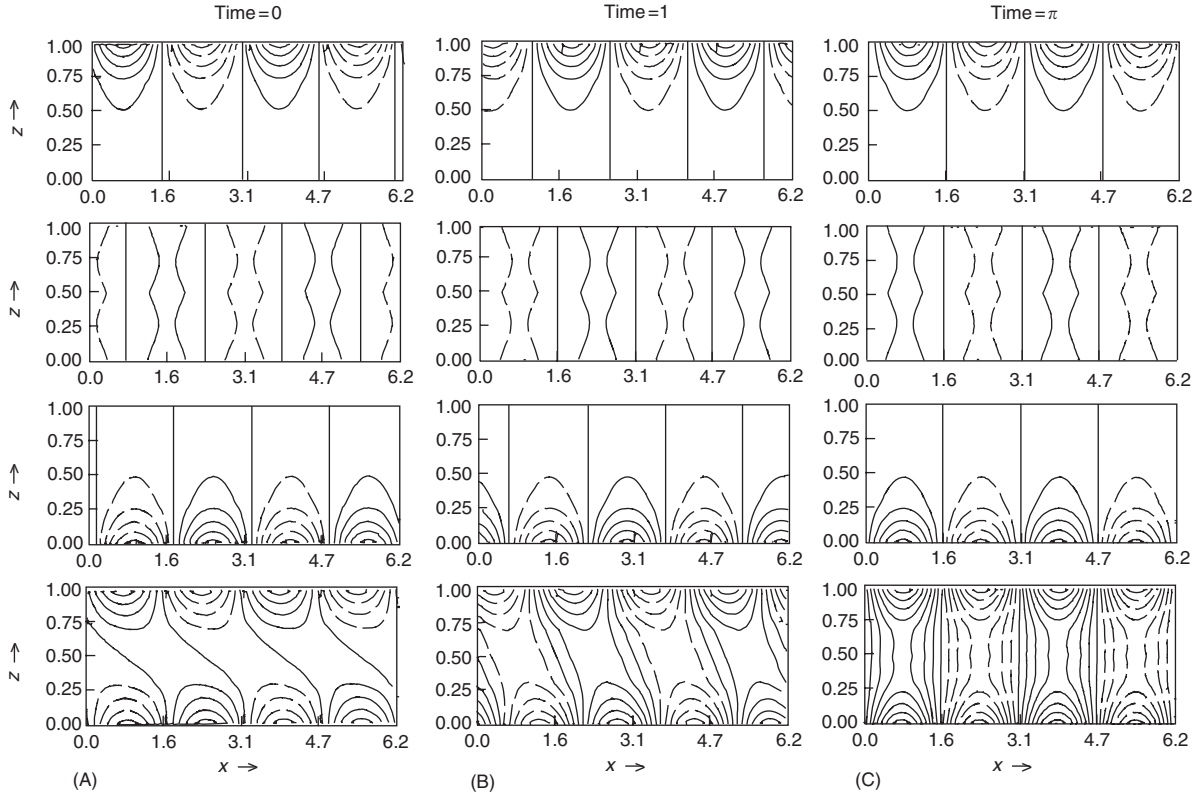
0076-P0175 Nonmodal growth can be quite strong in simple models like Eady's. However, most improvements to the model such as adding compressibility, variable Coriolis, and realistic vertical shear of U reduce nonmodal growth. Using more realistic initial states also tends to reduce nonmodal growth (e.g., using a wave packet instead of a wavetrain; using separate untilted upper and lower features instead of connecting them with a tilt).

Other Issues

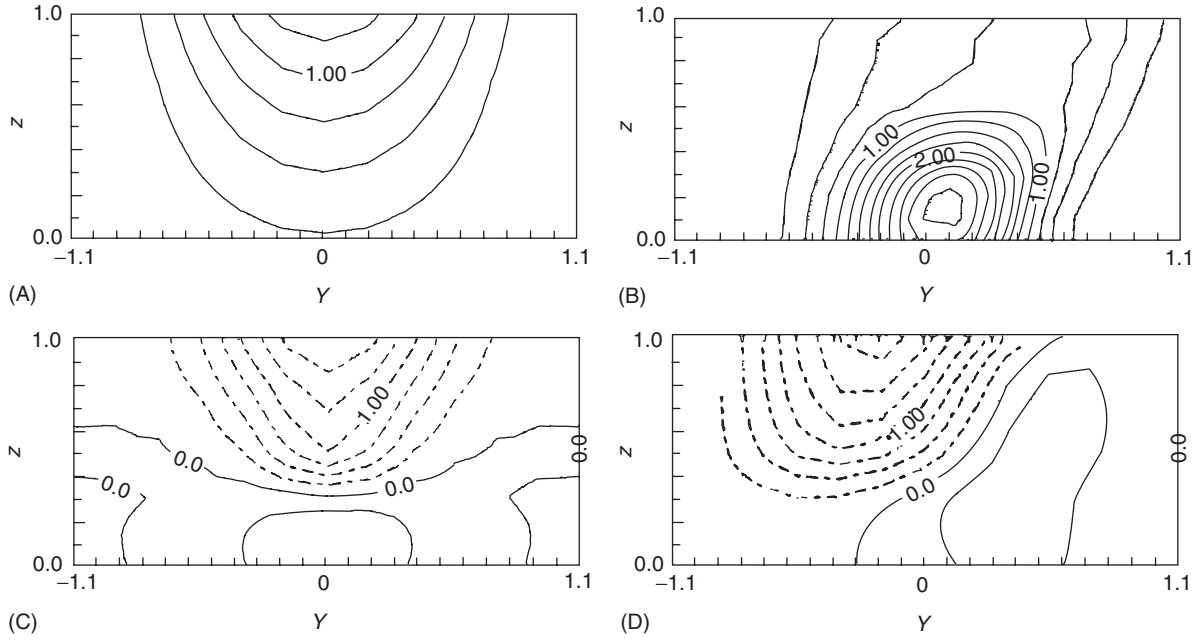
0076-P0180 Baroclinic instability has links with barotropic instability. First, each instability draws energy from mean flow shear. Second, barotropic instability has a similar stability criterion (absolute vorticity gradient changing sign in the domain). Third, there is interference between the two instabilities. The most unstable baroclinic eigenmode has optimal structure for a flow having only vertical shear, but when horizontal shear is added to that flow a different structure is needed or the eddy will be sheared apart. The subsequent structure is unlikely to be as optimal for baroclinic energy conversion. Hence, the baroclinic conversion will usually be reduced, though the barotropic growth mechanism may compensate. **Figure 7C** illustrates such a calculation; in this case adding a purely barotropic flow reduced the growth rate, even though the barotropic growth mechanism was activated.

0076-P0185 Baroclinically unstable frontal cyclones prefer to develop in certain regions. The preference may arise from local conditions, such as lower static stability or locally greater vertical shear. The illustrative model above assumes a wavetrain solution; when more localized development is considered, a variety of issues are raised.

0076-P0190 For example, if one uses a single low as the initial condition, the solution typically evolves into a chain of

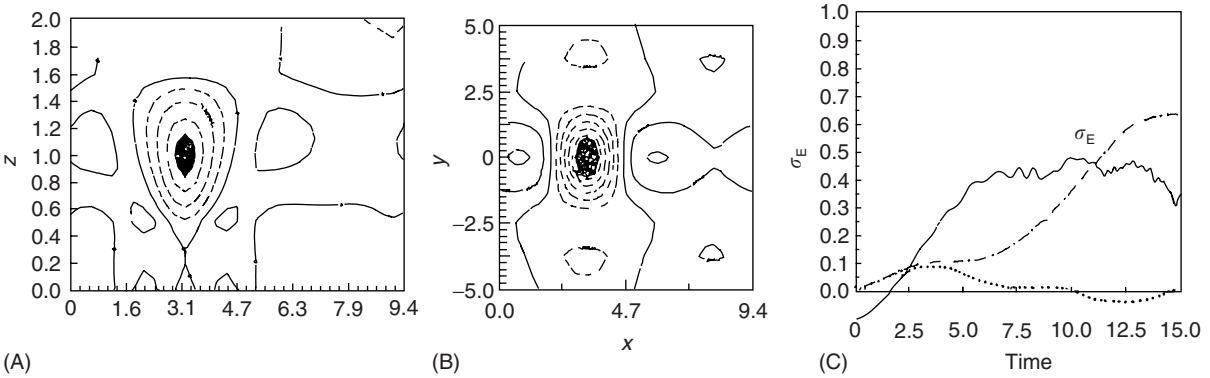


0076-F0006 **Figure 6** Nonmodal growth as a superposition process. Four initial value linear calculations are shown. The top three rows show three individual neutral continuum modes at three times. The bottom row shows the sum of the three modes at the initial time. (A) Initial condition; (B) time when energy growth is a maximum in the sum; (C) time when growth rate is zero in the sum. (Adapted with permission from Grotjahn R, Pedersen R and Tribbia J (1995) *Journal of Atmospheric Sciences* 36, 764–777).



0076-F0007 **Figure 7** Baroclinic energy conversion $(A_z \rightarrow A_e)$ for four models. (A) Lowest-order, square wave solution for an Eady-type model but including compressibility, increasing vertical shear in U, $\beta = 1$. (B) Solution when a surface frontal zone, centered at $Y = 0$, is added to the lowest-order mean flow U_0 and leading ageostrophic advective effects are included (using geostrophic coordinates). The frontal zone adds wind field: $0.2(2z - z^2)\ U_1$ where $U_1 = b_1(1 - \tanh^2(aY)ndash; b_2 - 3b_3 Y^2$ to U_0. The geostrophic coordinate transform causes the asymmetry. (C) Correction to the conversion shown in (A) when barotropically unstable horizontal shear U_1 is added to U_0. If the total wind is $U = U_0 + \mu U_1$, then the total conversion is $(a) + \mu(c)$. The barotropic shear reduces the growth rate. (D) Modification due to all leading order ageostrophic corrections. If those corrections are order μ, then the total conversion is $(a) + \mu(d)$. Ageostrophic conversions reduce the conversion and introduce asymmetry. (Adapted with permission from Grotjahn R (1979) *Journal of Atmospheric Science* 36, 2049–2074.)

waves as the modal constituents of the initial state disperse. Alternatively, a wave packet initial condition might be used consisting of a "carrier wave" multiplied by an amplitude envelope. The packet evolution depends upon the mean flow properties and assumptions made in the model. However, for reasonable choices of parameters, one might find a packet that spreads while propagating downwind. The leading edge of the packet has mainly faster, wider, and deeper modes. The trailing edge has slower, shorter, and



0076-F0008 **Figure 8** Initial value calculations for a linearly localized initial condition. (A) Zonal cross-section showing contours of streamfunction initially. Values < -1.0 are shaded. (B) Horizontal pattern of streamfunction at tropopause level $(z = 1.0)$ initially. The initial condition is constructed from neutral modes having similar phase speed. Growing or decaying normal modes are excluded. (C) Time-series of energy growth rate for three integrations. The linear model (dotted line) shows little growth since the nonmodal mechanism is weak and growing normal modes cannot develop. Also shown are nonlinear calculations for two amplitudes of the initial condition, where the solid line uses three times the initial amplitude of the dot-dashed line. Growing normal modes are activated by nonlinear interaction. Some evidence of nonlinear saturation is seen.

rwas.2002.0076 15/4/02 13:25 Ed:: M. SHANKAR No. of pages: 10 Pgn:: thilakam

shallower waves. It is possible to construct a localized structure that resists this dispersion by making a judicious combination of eigenmodes having similar phase speed, but different zonal wavenumber. **Figure 8** illustrates such an example using neutral continuum modes. When this model is solved as an initial value problem the packet maintains a localized shape for a long time. Almost no growth occurs because the normal modes were filtered out, and there is very slow phase shifting of the constituent modes. However, when nonlinear advection is allowed, modes interact and soon amplitude is injected into all the eigenmodes, including the growing normal modes, which grow rapidly in this example.

0076-P0195 Studies of regional development spawned subcategories of baroclinic instability. "Absolute" instability occurs when the wave packet expands faster than it propagates; the amplitude at a point keeps growing. "Convective" (in the advection sense) instability occurs when the packet moves fast enough so that growth then decay occurs as the packet moves past a point. "Global" instability (like the eigensolutions shown here) has growth that is invariant to a Galilean transform. Such is not the case for "locally" unstable modes. Normal modes for zonally varying basic states look like carrier waves modulated by a spatially fixed amplitude envelope; the envelope locally modifies the growth rate (sometimes called "temporal" instability). "Spatial" instability allows wavenumber to be complex while phase speed remains real.

0076-P0200 Nonlinear calculations raise other issues related to baroclinic instability. One issue concerns equilibration. The growing wave modifies the mean flow while drawing energy from it. This places a limit upon the cyclone development. In PV theory, this may be where the distortion shown in **Figure 5** becomes comparable to the cyclone width. Waves longer than the most unstable wave tend to reach a larger amplitude than the linearly most unstable mode. One reason for this is that they are deeper and so can potentially tap more APE in the mean flow. Another reason may be that the larger scale in both horizontal dimensions provides a longer time for PV contour distortion. Another possibility concerns the inversion of a PV anomaly: **Figure 4** uses the same magnitude of PV anomaly, but the streamfunction amplitude is larger for the broader PV anomaly.

0076-P0205 "Life cycle" studies model cyclones from birth to peak amplitude to decay. These studies typically find baroclinic growth followed by barotropic decay. This cycle fits the observed facts that eddies have a net heat flux and a net momentum convergence. These studies also reveal a characteristic evolution of the eddy structure: upper-level amplification compared to the linear eigenmodes. An explanation is that saturation is reached sooner at the critical level and at the surface while upper levels continue to grow. When averaged over the life cycle, the vertical distribution of the zonal mean eddy heat and momentum fluxes becomes more realistic.

Finally, the atmosphere has higher-order processes 0076-P0210
than the QG system. The biggest impact of ageostrophy is to break symmetries in the solutions. **Figure 7D** shows the leading order ageostrophic effects for a linear model. Ageostrophy causes enhanced eddy development on the poleward side (mainly by negative baroclinic conversion on the equatorward side), builds mean flow meridional shear, and slows down the wave. Ageostrophy also causes contours to be more closely spaced around a low and more widely spaced around a high.

See also

Barotropic Flow and Barotropic Instability (0077). **Cyclogenesis** (0129). **Cyclones:** Extra Tropical Cyclones (0128). **Dynamic Meteorology:** Balanced Flows and Potential-Vorticity Inversion (0140); Overview (0138); Waves and Instabilities (0141). **Fronts** (0039). **Quasi-geostrophic Theory** (0326). **Vorticity** (0449).

Further Reading

Gill A (1982) *Atmosphere–Ocean Dynamics*. New York: Academic Press.

Grotjahn R (1979) Cyclone development along weak thermal fronts. *Journal of Atmospheric Sciences* 336: 2048–2074.

Grotjahn R (1993) *Global Atmospheric Circulations: Observations and Theories*. New York: Oxford University Press.

Holton J (1992) *An Introduction to Dynamic Meteorology*, 3rd edn. San Diego, CA: Academic Press.

Hoskins B, McIntyre M, and Robertson A (1985) On the use and significance of isentropic potential vorticity maps. *Quarterly Journal of Royal Meteorological Society* 1111: 877–946.

James I (1994) *Introduction to Circulating Atmospheres*. Cambridge: Cambridge University Press.

Pedlosky J (1987) *Geophysical Fluid Dynamics*, 2nd edn. New York: Springer-Verlag.

Pierrehumbert R and Swanson K (1995) Baroclinic instability. *Annual Review of Fluid Mechanics* 227: 419–467.